Entropy in Causal Fermion Systems

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June 19, 2020, LQP Workshop

Preface

- Joint work with Felix Finster and Robert Jonsson
- Theory of causal fermion systems [3, p. 1]:
 - Novel approach to fundamental physics
 - Unification of interactions of the standard model with gravity
 - Close connections to quantum field theory
- Why consider entropy?
 - Plays important role in many physical effects
 - Many different definitions
 - Correlations not always well understood (e.g. area law)
 - Part of problems like black hole information paradox (see also [5] and [7])

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Outline



Preface

Preliminaries

- Causal Fermion Systems
- Causal Action Principle
- Surface Layer Integrals

Definition of the Entropy

- A Conserved Surface Laver Integral
- Definition of the Entropy

First Estimations

- A more specific Form of σ_{II}
- Minkowski Space Example
- Generalization to the Logarithm

Outlook

Causal Fermion Systems [3, p. 1, 3]

Definition (Causal Fermion System)

A causal fermion system (short CFS) of spin dimension $n \in \mathbb{N}$ ia a triple $(\mathcal{H}, \mathcal{F}, \rho)$ consisting of a complex separable Hilbert space $(\mathcal{H}, \langle . | . \rangle_{\mathcal{H}})$, the following subset of $L(\mathcal{H})$:

 $\mathcal{F} := \{ x \in L(\mathcal{H}) \mid x \text{ selfajoint with rank}(x) \le 2n \text{ and at most } n \text{ positive} \\ \text{and at most } n \text{ negative eigenvalues} \},$

and a measure ρ on \mathcal{F} , the so called *universal measure*.

Definition (Spacetime)

For a CFS $(\mathcal{H}, \mathcal{F}, \rho)$ we define *spacetime* as:

$$M := \operatorname{supp}(\rho) \subseteq \mathcal{F} \subseteq L(\mathcal{H})$$
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Causal Action Principle [3, p. 1-4]

Lagrangian: $\mathcal{L}: \mathcal{F} \times \mathcal{F} \to \mathbb{R}$, $(x, y) \mapsto \frac{1}{4n} \sum_{i,j=1}^{2n} \left(|\lambda_i^{xy}| - |\lambda_j^{xy}| \right)^2$,

with λ_i^{xy} , i = 1, ..., 2n (non-zero) eigenvalues of xy

- non-negative
- symmetric

Definition (Causal Action Principle)

Minimize the action

$$\mathcal{S}(
ho) := \int_{\mathcal{F}} \int_{\mathcal{F}} \mathcal{L}(x, y) d
ho(x) d
ho(y) \ ,$$

by varying ρ under the constraints

- Volume constraint: $\rho(\mathcal{F})$ constant,
- Trace constraint: $\int_{\mathcal{F}} \operatorname{tr}(x) d\rho(x)$ constant,
- Boundedness constraint: $\int_{\mathcal{F}} \int_{\mathcal{F}} \left(\sum_{i=1}^{2n} |\lambda_i^{xy}| \right)^2 d\rho(x) d\rho(y) \leq C$,

• From now on we always assume that ρ is such a minimizer

Surface Layer Integrals [3, p. 42-43]



Figure: Comparison of an ordinary surface integral and a surface layer integral [3, p. 42].

- In CFS no canonical measure on hypersurface
- \mathcal{L} decays fast for $|x y| > \delta$ (for Dirac particles: $\delta \sim m^{-1}$)

ightarrow gives integral over hypersurface smeared out on the scale of δ

A conserved surface layer integral [3, S. 42-45], [4]

- For $u \in \mathcal{H}$ arbitrary set $\mathcal{A}_u := \ket{u} \langle u \mid \in L(\mathcal{H})$
 - symmetric
 - finite rank
- \rightarrow One-parameter family of unitary transformations:

$$(U_{\tau})_{\tau \in \mathbb{R}} := (\exp(i\tau \mathcal{A}_{u}))_{\tau \in \mathbb{R}}$$

$$\Rightarrow \text{ Symmetry } \phi : \mathbb{R} \times \mathcal{F} \to \mathcal{F}, \ (\tau, x) \mapsto U_{\tau} x U_{\tau}^{-1} \text{ of } \mathcal{L}, \text{ i.e.:}$$

$$\mathcal{L}(\phi(\tau, x), y) = \mathcal{L}(x, \phi(-\tau, y)), \quad \forall \tau \in \mathbb{R}, x, y \in M$$

Lemma

For any $\Omega \subseteq M$ compact:

$$0 = \frac{d}{d\tau} \left(\int_{\Omega} d\rho(x) \int_{M \setminus \Omega} d\rho(y) \left(\mathcal{L}(\phi(\tau, x), y) - \mathcal{L}(x, \phi(\tau, y)) \right) \right) \Big|_{\tau=0}$$

=
$$\int_{\Omega} d\rho(x) \int_{M \setminus \Omega} d\rho(y) \left(D_{1, v_u(x)} - D_{2, v_u(y)} \right) \mathcal{L}(x, y) ,$$

with
$$v_u(x) := \frac{d}{d\tau} \phi(\tau, x)|_{\tau=0}$$
 for all $x \in \mathcal{F}$.

- $\Omega_{1/2}$: past of Cauchy surface $\partial \Omega_{1/2}$
- Ω_1 in the future of Ω_2
- Meantime Ω₂¹ := Ω₁ \ Ω₂ can be exhausted by compact sets (Ω̃_n)_{n∈ℕ}



 \Rightarrow For the integrand decaying fast enough at spatial infinity:

$$0 = \int_{\Omega_2^1} d\rho(x) \int_{M \setminus \Omega_2^1} d\rho(y) \Big(D_{1,v_u(x)} - D_{2,v_u(y)} \Big) \mathcal{L}(x,y) \\ = \Big(\int_{\Omega_1} \int_{M \setminus \Omega_1} - \int_{\Omega_2} \int_{M \setminus \Omega_2} \Big) \Big(D_{1,v_u(x)} - D_{2,v_u(y)} \Big) \mathcal{L}(x,y) d\rho(y) d\rho(x)$$

$$\rightarrow \langle u, u \rangle_{\rho} := \int_{\Omega} d\rho(x) \int_{M \setminus \Omega} d\rho(y) \Big(D_{1, v_{u}(x)} - D_{2, v_{u}(y)} \Big) \mathcal{L}(x, y)$$
conserved for Ω past of Cauchy surface $\partial \Omega$

Definition of the Entropy

• In Minkowski space limiting case (see [4, Chapter 5]):

 $\langle u, u \rangle_{\rho} = c \langle u | u \rangle_{\mathcal{H}} \quad \forall u \in \mathcal{H}_{F} \subseteq \mathcal{H} \text{ closed subspace }.$

- \rightarrow Assume this holds in general for \mathcal{H}_F suitable
- $\rightarrow \langle .,. \rangle_{
 ho}$ inner product on \mathcal{H}_{F}
 - Spatial localization for $U \subseteq M$:



$$\underbrace{\frac{1}{2} \Big(\int_{\Omega \cap U} \int_{M \setminus \Omega} + \int_{\Omega} \int_{(M \setminus \Omega) \cap U} \Big) \Big(D_{1, v_u(x)} - D_{2, v_u(y)} \Big) \mathcal{L}(x, y) d\rho(x) d\rho(y)}_{=: \langle u | u \rangle_{U, \rho}}$$

Define localized state operator

$$\langle u|u\rangle_{U,\rho} = \langle u, \sigma_U u\rangle_{\rho}, \quad \forall u \in \mathcal{H}_F$$

Von-Neumann-like entropy (for reduced one-particle density) [6, (34)], [8, p. 400-401]

$$S := \operatorname{tr}_{\mathcal{H}_F} (\sigma_U \log(\sigma_U) + (1 - \sigma_U) \log(1 - \sigma_U))$$

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A more specific Form of σ_U [3, p. 39-40], [4, chapter 5]

• Using a kernel *Q* the product $\langle ., . \rangle_{\rho}$ can be rewritten as:

$$\langle u|u\rangle_{
ho} = 4 \mathrm{Im} \int_{\Omega} d\rho(x) \int_{M\setminus\Omega} d\rho(y) \prec \psi^{u}(x), Q(x,y)\psi^{u}(y) \succ_{y}$$

And correspondingly:

$$\langle u|u\rangle_{U,
ho} = 2\mathrm{Im}\Big(\int_{\Omega} d
ho(x) \int_{M\setminus\Omega} d
ho(y) \prec \psi^{u}(x), Q(x,y)\chi_{U}(y)\psi^{u}(y) \succ_{y} + \prec \chi_{U}(x)\psi^{u}(x), Q(x,y)\psi^{u}(y) \succ_{y}\Big)$$

 \rightarrow More specific form for σ_U :

$$\sigma_U = \frac{1}{2}((\chi_U)^* + \chi_U) ,$$

with $(\chi_U)^*$ adjoint of χ_U wrt $\langle . | . \rangle_{\rho}$

• In vakuum Mionkowski space: $\sigma_U = \sigma \tilde{\sigma}_U \sigma$ with

• $\sigma: L^2(\mathbb{R}^3, \mathbb{C}^4) \to \mathcal{H}_F$ projection operator with kernel $\sigma(\vec{x}, \vec{y})$ s.t.:

$$\sigma(ec{x},ec{z}) = \int_{\mathbb{R}^3} d^3ec{y} \sigma(ec{x},ec{y}) \sigma(ec{y},ec{z}) \; .$$

• $\tilde{\sigma}_U: L^2(\mathbb{R}^3, \mathbb{C}^4) \to L^2(\mathbb{R}^3, \mathbb{C}^4)$ integral operator with kernel

$$\tilde{\sigma}_U(\vec{x},\vec{y}) = \frac{1}{2} \big(\chi_U(\vec{x}) + \chi_U(\vec{y}) \big) \sigma(\vec{x},\vec{y}) \; .$$

 \Rightarrow Assume in the following that σ_U can always be represented in this form

Consider $\sigma_U - (\sigma_U)^2$

- $\log(\sigma_U)$ difficult to calculate
- Why log-formula at all?
- $x x^2$ similar behaviour:
 - EVs 0, 1 don't contribute
 Measues mixing

plus easier to calculate (similar as in [2, Theorem 3.1]):



$$\begin{aligned} \operatorname{tr}(\sigma_{U} - (\sigma_{U})^{2}) &= \operatorname{tr}(\sigma_{U} - \sigma_{U}(\sigma - \sigma_{\mathbb{R}^{3}\setminus U})) = \operatorname{tr}(\sigma_{U}\sigma_{\mathbb{R}^{3}\setminus U}) = \operatorname{tr}(\tilde{\sigma}_{U}\sigma\tilde{\sigma}_{\mathbb{R}^{3}\setminus U}\sigma) \\ &= \frac{1}{4} \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}} \int_{\mathbb{R}^{3}} \left(\chi_{U}(\vec{x}_{1}) + \chi_{U}(\vec{x}_{2}) \right) \left(\chi_{\mathbb{R}^{3}\setminus U}(\vec{x}_{3}) + \chi_{\mathbb{R}^{3}\setminus U}(\vec{x}_{4}) \right) \cdot \\ &\qquad \operatorname{tr}_{\mathbb{C}^{4}} \left(\sigma(\vec{x}_{1}, \vec{x}_{2}) \sigma(\vec{x}_{2}, \vec{x}_{3}) \sigma(\vec{x}_{3}, \vec{x}_{4}) \sigma(\vec{x}_{4}, \vec{x}_{1}) \right) d^{3}\vec{x}_{1} d^{3}\vec{x}_{2} d^{3}\vec{x}_{3} d^{3}\vec{x}_{4} \\ &= \int_{U} d^{3}\vec{x} \int_{\mathbb{R}^{3}\setminus U} d^{3}\vec{y} \operatorname{tr}_{\mathbb{C}^{4}} \left(\sigma(\vec{x}, \vec{y}) \sigma(\vec{y}, \vec{x}) \right) , \end{aligned}$$

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Minkowski Space Example (see also [3, p. 25-34])

 Vacuum Minkowski space: σ = Π^ε regularized projector to negative frequency solutions of Dirac equation



→ Split surface layer integral with cutoff constant 0 < $\varepsilon \ll \delta \ll I$:

$$\int_{B_l(0)}\int_{\mathbb{R}^3\setminus B_l(0)}\cdots=\int_{B_l(0)}\int_{B_l(0)^c\cap B_\delta(\vec{x})}\cdots+\int_{B_l(0)}\int_{B_l(0)^c\cap B_\delta(\vec{x})^c}\ldots$$

 In the limit ε → 0 (physical case) fist term diverges, second one bounded → focus on first term • Simplifying integrals, the first term becomes:

$$\int_{B_{l}(0)} d^{3}\vec{x} \int_{B_{l}(0)^{c} \cap B_{\delta}(\vec{x})} d^{3}\vec{y} \operatorname{tr}_{\mathbb{C}^{4}}(\Pi^{\varepsilon}_{-}(\vec{x},\vec{y})\Pi^{\varepsilon}_{-}(\vec{y},\vec{x}))$$
$$= C \cdot \left[l^{2} \int_{0}^{\delta} dr \, r^{3} F_{\varepsilon}(r^{2}) - \frac{1}{12} \int_{0}^{\delta} dr \, r^{5} F_{\varepsilon}(r^{2}) \right],$$

with $F_{\varepsilon} : \mathbb{R} \to \mathbb{C}$ such that $F_{\varepsilon}(r^2) = \operatorname{tr}_{\mathbb{C}^4}(\Pi^{\varepsilon}_{-}(0, r\hat{e}_1)\Pi^{\varepsilon}_{-}(0, r\hat{e}_1))$ for all $r \ge 0$

- First integral dominates for $I \gg \delta$
- \rightarrow obtain area-law-like form:

$$\operatorname{tr}(\sigma_{B_{l}(0)} - \sigma_{B_{l}(0)}^{2}) = \underbrace{\mathcal{C}(\varepsilon)}_{\varepsilon \to 0} \cdot \left(\mathcal{C}_{0}(\delta) \cdot l^{2} + \underbrace{\mathcal{C}_{1}(\delta)}_{\ll \mathcal{C}_{0}(\delta) l^{2} \text{ for } l \gg \delta} \right) + \underbrace{\mathcal{D}(\varepsilon)}_{\mathcal{O}(1) \text{ for } \varepsilon \to 0} \cdot \operatorname{Corrections},$$

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Generalization to the logarithm

• Taylor expansion of logarithm:

$$\operatorname{tr}\left((1-\sigma_U)\ln(1-\sigma_U)\right) = -\sum_{n=1}^{\infty} \frac{(-1)^n}{n} \operatorname{tr}(\sigma_U^n) + \sum_{n=1}^{\infty} \frac{(-1)^n}{n} \operatorname{tr}(\sigma_U^{n+1}) ,$$

 \rightarrow Additivity:

$$\operatorname{tr}(\sigma_U^{n+1}) = \operatorname{tr}((\sigma_U)^n (\sigma - \sigma_{\mathbb{R}^3 \setminus U})) = \operatorname{tr}((\sigma_U)^n \sigma) - \operatorname{tr}((\sigma_U)^n \sigma_{\mathbb{R}^3 \setminus U}),$$

- \rightarrow By definition of σ_U : tr($(\sigma_U)^n \sigma$) = tr($(\sigma_U)^n$)
- \rightarrow And similar as before: tr $((\sigma_U)^n \sigma_{\mathbb{R}^3 \setminus U}) = tr((\chi_U \sigma)^n (\chi_{\mathbb{R}^3 \setminus U} \sigma))$
- \rightarrow Finially iteratively use:

$$\operatorname{tr}\left((\chi_U\sigma)^n(\chi_{\mathbb{R}^3\backslash U}\sigma)\right) = \operatorname{tr}\left((\chi_U\sigma)^{n-2}(1-\chi_{\mathbb{R}^3\backslash U})\sigma(\chi_U\sigma)(\chi_{\mathbb{R}^3\backslash U}\sigma)\right)$$
$$=\operatorname{tr}\left((\chi_U\sigma)^{n-1}(\chi_{\mathbb{R}^3\backslash U}\sigma)\right) - \operatorname{tr}\left((\chi_U\sigma)^{n-2}(\chi_{\mathbb{R}^3\backslash U}\sigma)(\chi_U\sigma)(\chi_{\mathbb{R}^3\backslash U}\sigma)\right),$$

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- Proceed similarly with $tr(\sigma_U \log(\sigma_U)) = tr(\sigma_U \log(1 \sigma_{\mathbb{R}^3 \setminus U}))$
- $\rightarrow\,$ After simplification this yields expansion:

$$\operatorname{tr}\left(\sigma_U \log(\sigma_U) + (1 - \sigma_U) \ln(1 - \sigma_U)\right) = \sum_{n=1}^{\infty} \alpha_n \operatorname{tr}\left(\left((\chi_U \sigma)(\chi_{\mathbb{R}^3 \setminus U} \sigma)\right)^n\right)$$

- Resulting traces correspond to higher order surface layer integrals (alternating integrals over inside and outside)
- Conjecture:
 - Short range of integrand → for each pair of integrals over U and U^c only small strips of space around ∂U relevant
 - Surface layer integrals decay with growing order
 - Only lower orders in n relevant
 - Lowest order: Ordinary surface layer integral, which likely has area-law-like form

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Outlook

• Prove more general area-law-like form for area A_U of U:

$$S = \underbrace{C(\varepsilon)}_{\varepsilon \to 0} \cdot \left(C_0(\delta) \cdot A_U + \underbrace{C_1(\delta)}_{\ll C_0(\delta)A_U \text{ for } A_U \gg \delta^2} \right) + \underbrace{D(\varepsilon)}_{\mathcal{O}(1) \text{ for } \varepsilon \to 0} \cdot \text{Corrections}$$

Show that S := tr(σ_U log(σ_U) + (1 − σ_U) log(1 − σ_U)) really corresponds to entanglement entropy

 \rightarrow Consequences for information paradox: General problem [5], [7]:

- Independent of the initial state a black hole evaporates by emitting Hawking radiation
- ► Shrinking horizon → Bekenstein-Hawking entropy shrinks
- Thermal radiation \rightarrow Entanglement-entropy grows
- Problem if Bekenstein-Hawking entopy and entanglement-entropy are always equal
- Hope: Answer question whether they really are equal all the time



[1]: CFS-website: https://causal-fermion-system.com/



[2]: Chris Brislawn: *Traceable Integral Kernels on Countably Generated Measure Spaces*, Pacific Journal of Mathematics, **150** (1991), no. 2.



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[5]: Daniel Harlow: *Jerusalem Lectures on Black Holes and Quantum Information*, arXiv:1409.1231 [hep-th], Rev. Mod. Phys., **88** (2016), no. 1 (version 4, 2015).



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[8]: Walter Thirring: *Quantum Mathematical Physics: Atoms, Molecules and Large Systems*, Springer, Berlin, Heidelberg 2002 (second edition, second printing 2003).

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